

Exploring the Universe, Sharing its Wonders

# Lessons learned from Gemini observations of Didymos mutual events in 2017



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## Many ways to get time for the many ways we do research!



	Regular Programs	Large & Long Programs	Fast Turnaround	Director's Discretionary Time	Poor Weather
When	Once per semester	Once per <b>year</b>	Once per month	Anytime	Anytime
Review	National Allocation Committees	LLP Time Allocation Committee	Other proposers! (No TAC)	Chief Scientist	Head of Science Operation
What for?	Routine programs	Large allocation and/or multi- semester	Immediate, short and/or follow-up	Special opportunities	Bright targets
Observing mode	Queue, Classical, Priority Visitor	Priority Visitor (Queue occasionally)	Queue	Queue, Priority Visitor	Queue

65% 20% 10% 5% Weather loss

## "regular" projects

- Submitted during the Calls for Proposals (CfP):
   in October (A semester) and March (B semester)
- Programs lifetime is 6 months:
   Feb-Jul for A semester and Aug-Jan for B semester
- Full access to all capabilities
- Classical or Queue
- Oversubscription ~2 for US



## LLP: projects that require 100s of hours and/or several semesters

- One proposal and you're done except for progress reports
- Letter of intent required in February, and proposal submitted with the B semester CfP
- Can stay active for up to 3 years!
- Oversubscription ~5 at first but has dropped to ~2

24 publications so far since 2015 on:

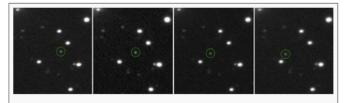


Figure 6. Sequence of four Gemini images of near-Earth asteroid 2014 EN45 (circled). This object was discovered by the NEOWISE survey on 6 March 2014 and imaged by GMOS-S on 13 March 2014, which provided critical astrometry needed to confirm the asteroid's orbit. 2014 EN45 is ~800 meters in diameter and is as dark as a piece of coal.

Exoplanets Lensed galaxies

Asteroids Low-mass galaxy evolution

Kuiper Belt objects Faint Halo objects

Supernovae Quasars

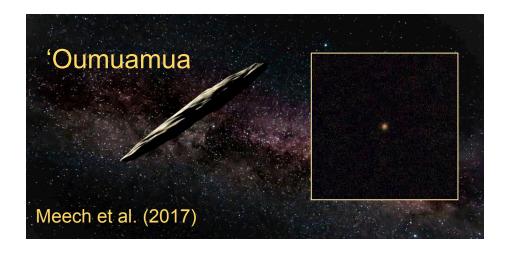
### FT: rapid response

- Submitted every month (e.g., 26 Feb)
  - O Evaluation starts the first of the following month (e.g., 1 Mar)
  - Answer sent the 3rd week (e.g., 15 Mar)
  - O Program starts the next month (e.g., 1 Apr)
  - Program ends after 3 month (e.g., 30 Jun)
- All proposers serve as referee
- Limited access to Visiting Instruments
- Oversubscription varies
   37 publications so far since 2015 on various topics!



## **DDT:** extraordinary opportunities

- Submitted directly to the Chief Scientist (John Blakeslee)
- High priority programs
- Report on work progress 4 months after observations
- Oversubscription "auto-regulated"



### PW: when we open on cloudy nights

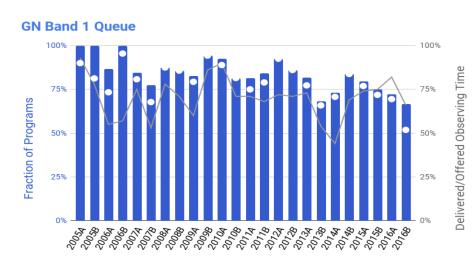
- Submitted directly to the Head of Science Operation no TAC
- For bright targets (typical 2-4m telescope projects)
- For poor seeing (1") and/or thick clouds (extinc. > 1mag)
- Approved based on schedule
- Best effort basis



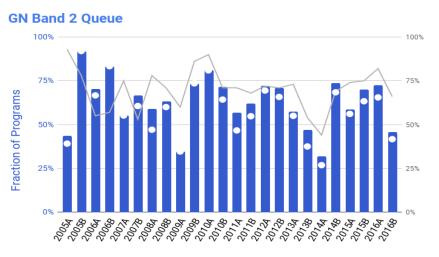
# The likelihood to get data (in queue mode) depends on:

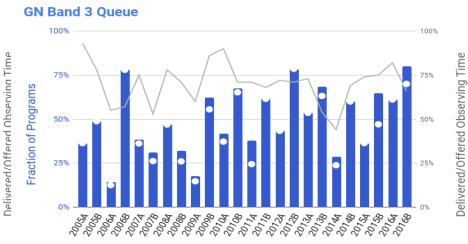
- 1. Program Band (1, 2, 3)
- 2. Program type (Q, LP, DD, FT, PW)
- 3. Program restrictions (timing windows, weather, frequency)
- 4. Program status (ToO, started, intru. configuration, ...)

## The likelihood to get data depends on 1. the Band



- Band 1 has absolute priority
- Band 3 are "fillers"
- Band 2 are successful when the semester goes well





# The likelihood to get data depends on 2. the program type

(Priority increases from bottom to top)

DD	FT	Q	LP	Others
Band 1	<b>→</b>			
	Band 1	Band 1	Band 1	
Band 2	<u>+</u>			
	Band 2	Band 2	Band 2	ENG
		Band 3		
				Band 4 (PW)

NOTE: this is a methodology

# The likelihood to get data depends on 3. the program restrictions

- **Timing Windows**: Timing windows matter, and observations can be scheduled in better conditions than requested if needed to try and meet a timing window.
- Time frequency: Same as for Timing Windows, time frequency requirement is followed (after the priority based on the Bands has been considered).
- Weather constraints:
  - Best weather conditions (IQ20%, CC50%) happen 10-15% of the time.
  - Execution in better than requested conditions will on average lead to lower completion rates!
  - Water Vapor and Sky Background restrictions are taken into account.

NOTE: this is a methodology

# The likelihood to get data depends on 4. the program status

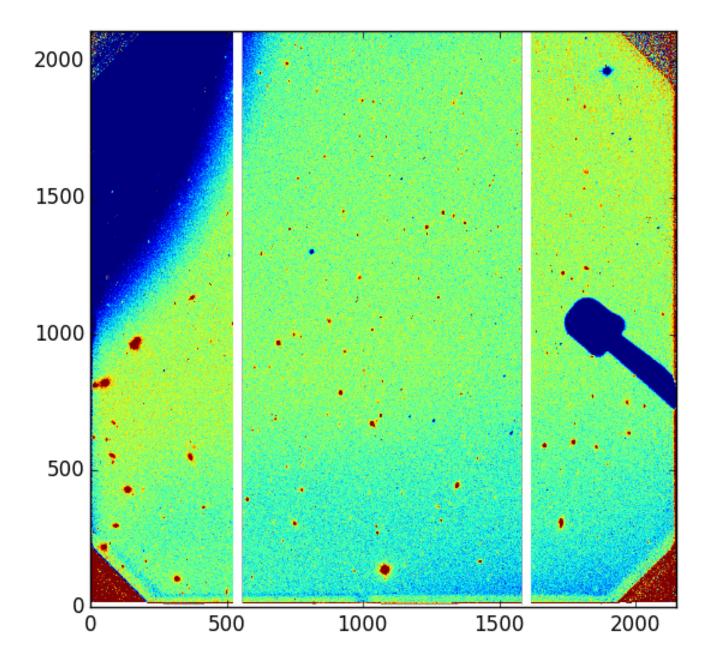
- Rapid ToOs: Take absolute priority (within their band) when they are triggered.
- Standard ToOs: Need to be done before their timing window ends (days to weeks).
- **Current completion:** Programs that have been started should get priority over non-started ones, until they reach 80% completion.
- **R.A.**: Earlier targets take priority if we are about to lose them.
- Instrument configuration: Between two equivalent programs, one can take priority
  if it needs the same configuration as a higher priority program.
- **Visiting instruments programs**: Because they are "block scheduled", they get higher priority when the instrument is on the telescope.
- Known thesis projects: They are completed first

NOTE: this is a methodology

### Band 3 program are "fillers"

#### Will only be started if:

- Weather conditions not covered by Band 1 and 2
- RAs not covered by Band 1 and 2
- Has a reasonable chance of getting completed, or a useful subset of observations can be completed.
- Brownie points for thesis project.
- If it is a GMOS program, it is better the configuration is commonly used by Band 1 and 2 programs.

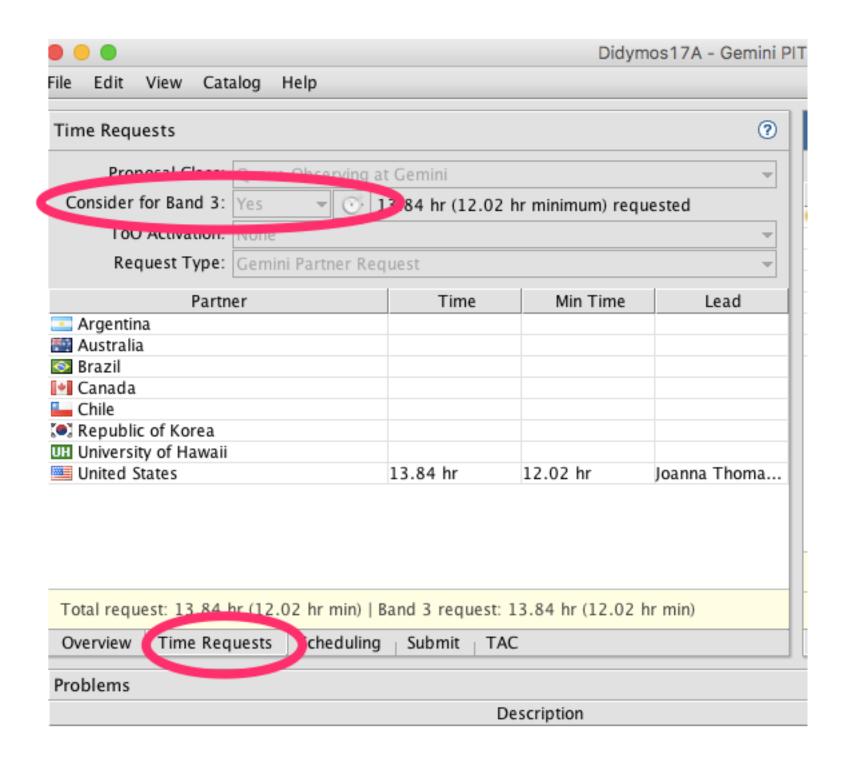


## US UCG representatives

- Karen Meech (Chair and UH rep but the one who brought this up)
- Matt Bayliss (CfA)
- Jessica Werk (UW)
- Vini Placcus (Notre Dame)
- Mark Brodwin (U. Missouri)
- http://www.gemini.edu/science/#ucg

## NOAO User's Committee

http://ast.noao.edu/about/committees/usercom



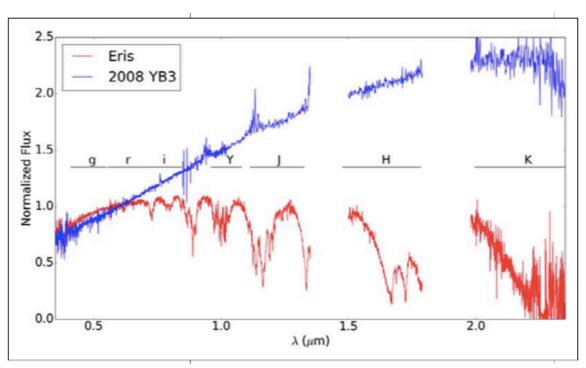
#### **Lessons Learned**

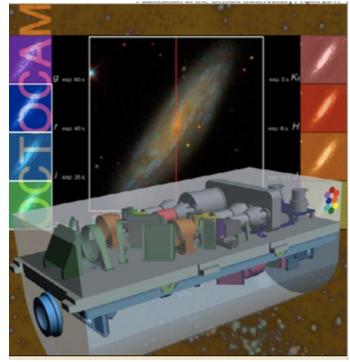
- Lesson #1: Be prepared submit FT proposals shortly after the regular program notifications (June and December)
- Lesson #2: Relax sky brightness condition as a last resort. For solar system objects sky brightness can be a double whammy because of the likelihood of getting close to the waxing or waning moon.
- Lesson #3: Until GMOS bubbles are corrected, track non-sidereally or dither to keep target in the central area of CCD2.
- Lesson #4: US Solar System Community needs to advocate for better representation amongst the rankings that come out of the combination of the sub-TACs.
- Lesson #5: If you need timing windows, say no to band 3.



## formerly OCTOCAM

8-channel imager and spectrograph capable of multi-band imaging, long slit broad-band spectroscopy and high-time-resolution (Gemini-South)





Contract Signed on March 3, 2017 CoDR passed on August 3, 2017 Preliminary Design Review April 4-5, 2018 Expected Commissioning March 2022









